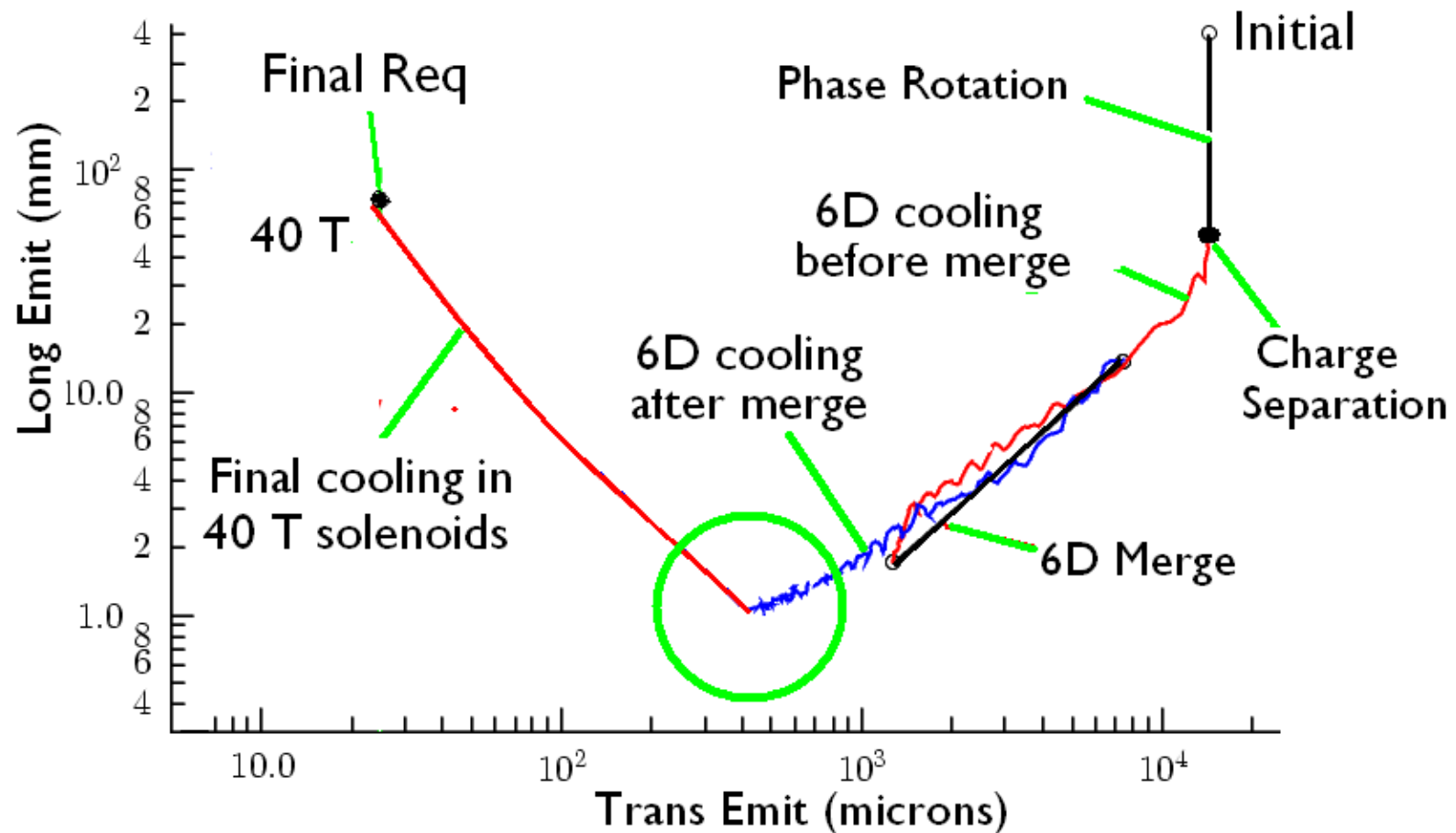




# Space charge effects in 6D cooling

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MAP Winter Meeting  
SLAC



This is a plot from last year's Winter Collaboration Meeting in JLAB

# Outline

- Introduction
- Transverse Space-Charge
- Longitudinal Space-Charge
- The Challenge: We should avoid cooling to low longitudinal emittance
- Solution
  1. New taper after merge with less emittance exchange
  2. Added new cooling lattice for lower transverse emittance
- Improvements if HTS used in 6D
- Conclusion

# 1) Introduction

- For  $10^{34} \text{ s}^{-1} \text{ cm}^{-2}$  at 1.5 TeV, we need  $2 \times 10^{12}$  muons/bunch
- At end of 'Final Cooling' the bunch is long (4 m) giving a relatively moderate current and space charge effects
- But in late 6D cooling, the rms bunch length is only  $\sigma_{ct} \approx 2 \text{ cm}$
- And when losses in 'Final Cooling' and Acceleration are included the bunches have  $4.87 \times 10^{12}$  muons, giving a peak current of

$$\frac{4.87 \times 10^{12} \times 1.6 \times 10^{-19} \times 3 \times 10^8}{\sqrt{2\pi} \times 0.02} = 4.6 \text{ kA} \quad \text{at only } 160 \text{ MeV/c}$$

- This is a large current, at a low mom, so coherent effects can be expected

# Transverse Space Charge Tune Shift

For a Gaussian longitudinal charge distribution with rms length of  $\sigma_z$  the tune shift per unit length is

$$\frac{d\nu_{\text{space}}}{dL} = \left( \frac{N_\mu}{\sqrt{2\pi} \sigma_z} \right) \frac{r_\mu}{4\pi \epsilon_\perp \beta_v \gamma^2}$$

For a cell of a periodic system,  $\nu = \Delta\phi/L_{\text{cell}}$  the fractional tune shift is thus

$$\frac{\Delta\nu_{\text{space}}}{\nu_{\text{cell}}} = \left( \frac{N_\mu}{\sqrt{2\pi} \sigma_z} \right) \frac{r_\mu L_{\text{cell}}}{\phi_{\text{cell}} 4\pi \epsilon_\perp \beta_v \gamma^2}$$

# Transverse tune shift in late 6D cooling

The phase advance per cell, as a function of momentum, in an RFQFO Guggenheim lattice is from  $\pi$  to  $2\pi$ . So the average phase advance per cell is  $1.5 \pi$  and  $\Delta\nu = \pm\pi/2$

$L_{cell} = 0.6875$  m long,  
momentum =  $207 \text{ MeV}/c$ ,  
 $N_{\mu} = 4.87 \cdot 10^{12}$ ,  
 $\sigma_{ct} = 1.87$  cm, and  
 $\epsilon_{\perp} = 400 \mu\text{m}$

Giving  $\Delta\nu/\nu = 0.068$

a small fraction of the acceptance of  $\pm 0.25/0.75 = 0.33$

So this is not a problem

# Longitudinal Space Charge

Published calculations of longitudinal space charge are mostly given for uniform charge densities within 3 dimensional ellipsoids whose projections in all coordinates are parabolic:

$$\frac{dQ}{dx} = \rho_o \left( 1 - \frac{z^2}{z_m^2} \right)$$

For these distributions the rate of change of the longitudinal space charge field, i.e the longitudinal defocusing force is

$$\mathcal{E}'_{sc} = \frac{d\mathcal{E}'_{sc}}{dz_{lab}} = \left( \frac{3 Q g}{8 \pi \epsilon_o \gamma^2 (z_m)^3} \right)$$

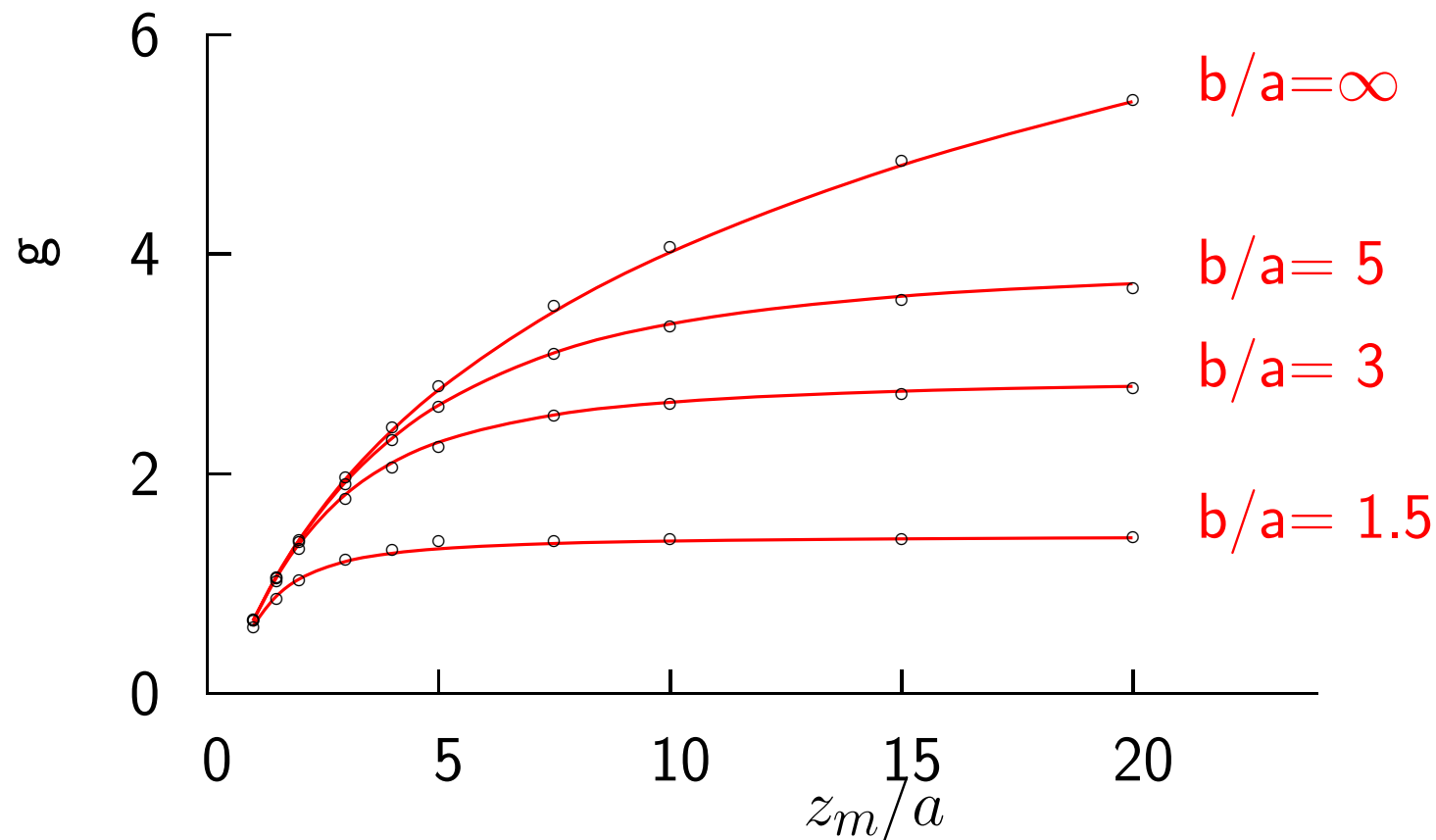
where  $g$  is a form factor that depends on  $z_m/a$  and  $b/a$  where  $z_m$  is the longitudinal, and  $a$  is the transverse size of the ellipsoid, and  $b$  is the assumed pipe radius

Note inverse cubic dependence on bunch length

## Form factor $g$

In the absence of a pipe, the space charge, as given above, is linear in  $z$  with a fixed slope  $\mathcal{E}'$  so that  $g$  is a constant for any given  $z_m/a$

In a pipe it deviates from linearity, but average values of  $g$  have been numerically calculated (Allen) and shown below with a fit



## rf Requirement

To reach the same performance as simulated without space charge, the needed  $\mathcal{E}'_{\text{rf}}$  must equal the simulated contribution  $\mathcal{E}'_{\text{sim rf}}$  plus the magnitude of the space charge defocus  $\mathcal{E}'_{\text{sc}}$

$$\frac{\mathcal{E}'_{\text{rf}}}{\mathcal{E}'_{\text{sim rf}}} = 1 + \xi \quad \text{where} \quad \xi = \frac{\mathcal{E}'_{\text{sc}}}{\mathcal{E}'_{\text{sim rf}}}$$

Assuming bunch length small compared with the rf wavelength:

$$\mathcal{E}'_{\text{sim rf}} = \frac{\omega \mathcal{E}_{\text{rf}} \eta}{c} \cos(\phi)$$

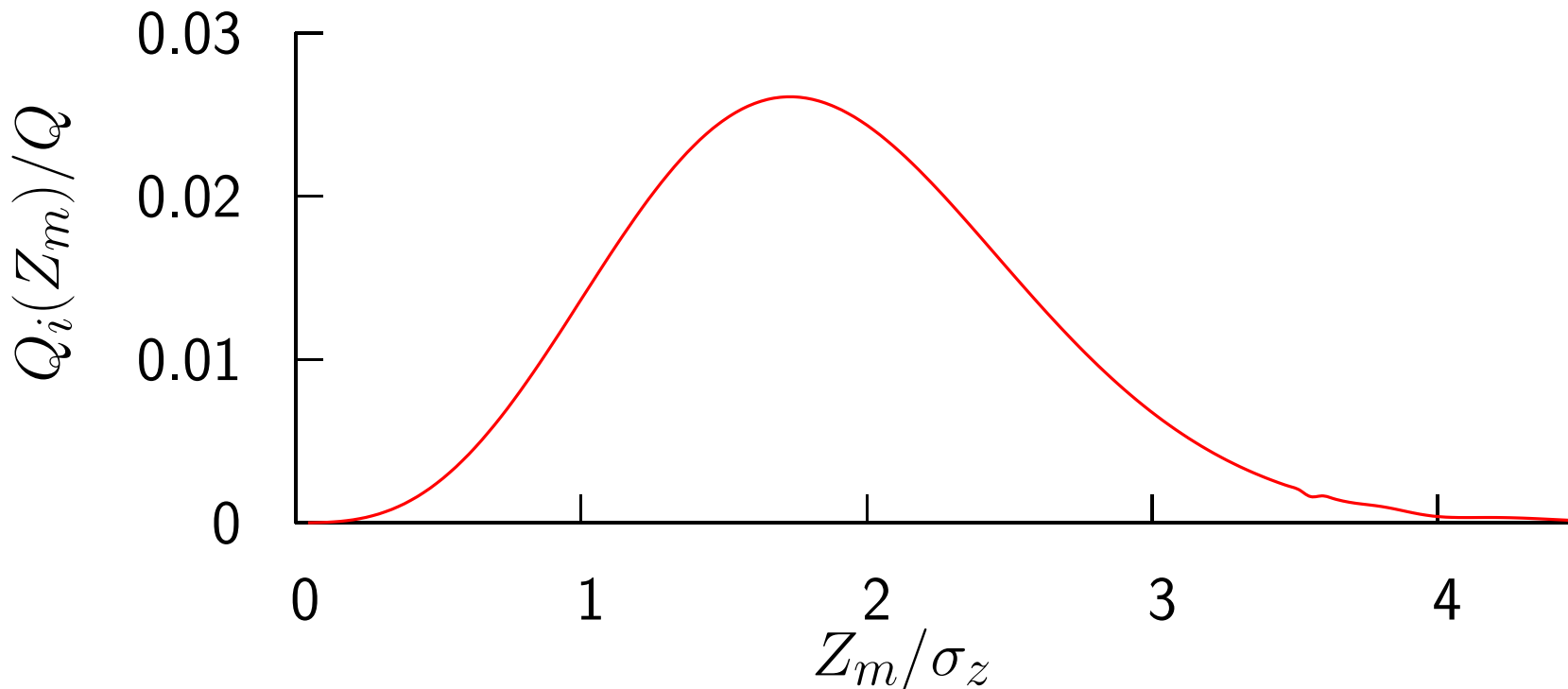
where  $\phi$  is the rf phase with respect to the zero crossing,  $\omega$  is the rf frequency, and  $\mathcal{E}_{\text{rf}}$  is the simulated rf gradient, and  $\eta$  is the fraction of the lattice filled with rf.

$$\xi = \frac{c}{\omega \mathcal{E}_{\text{rf}} \eta \cos(\phi)} \mathcal{E}'_{\text{sc}}$$



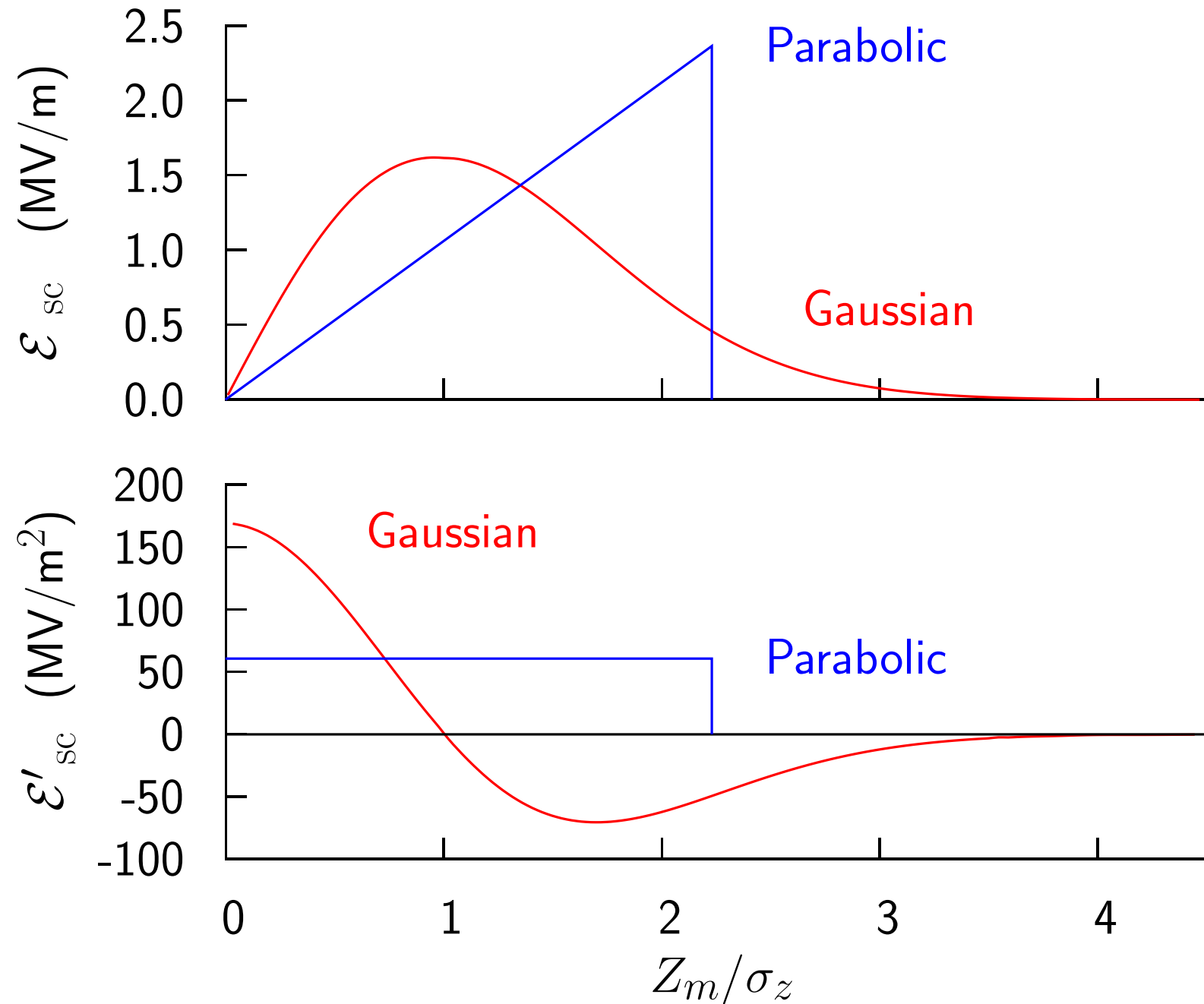
## A Gaussian by sum of Parabolics

Since all the above calculations are linear, we can form a sum of parabolic distributions of charges  $Q_i$  and lengths  $z_m$  that sum to a Gaussian distribution approximating those simulated for cooling in the absence of space charge. For 100 such distributions:

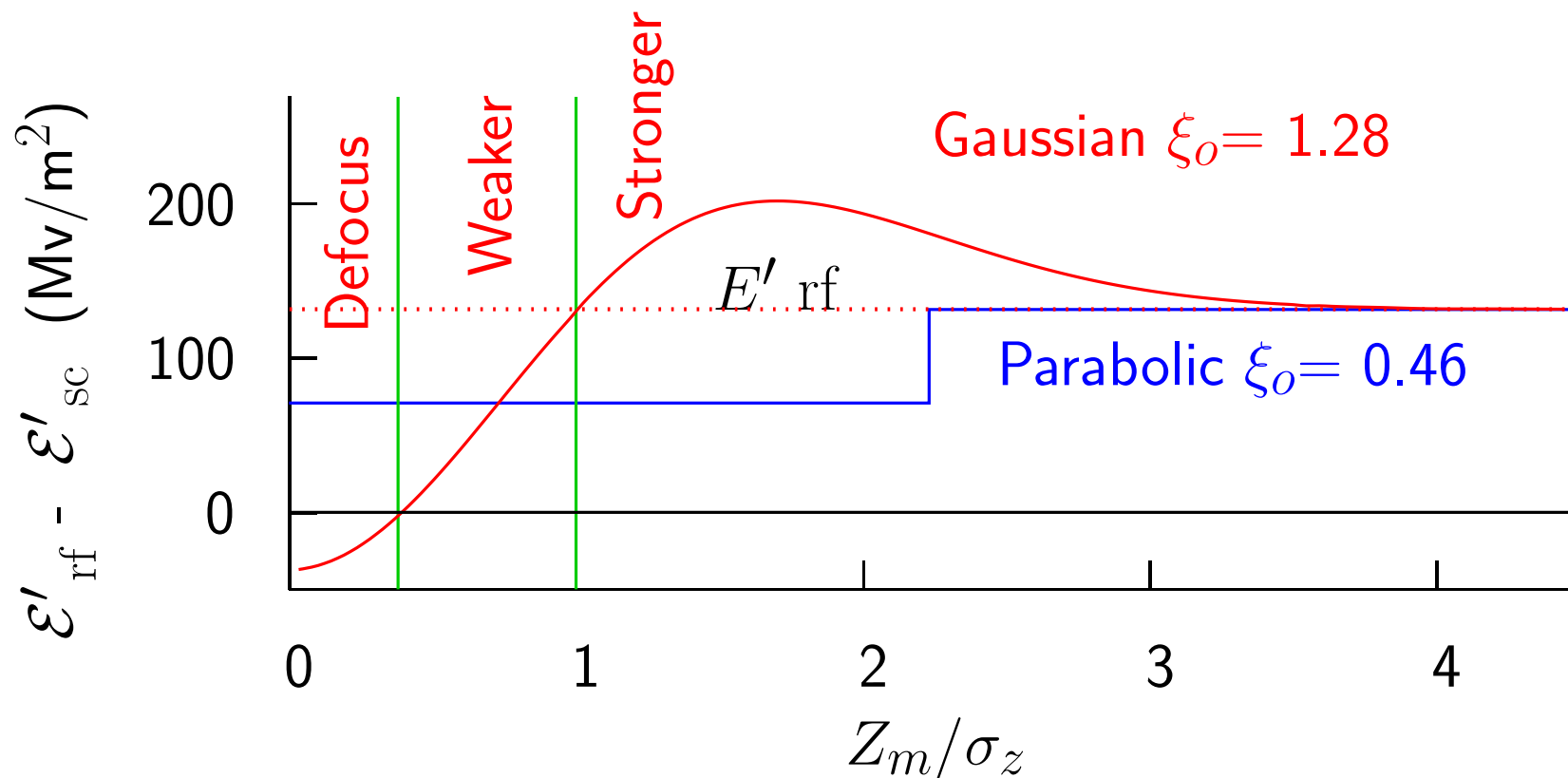


In the following we show calculations for both a Gaussian and Parabolic distributions.

# Space Charge $\mathcal{E}$ and $\mathcal{E}'$ for JLAB scheme



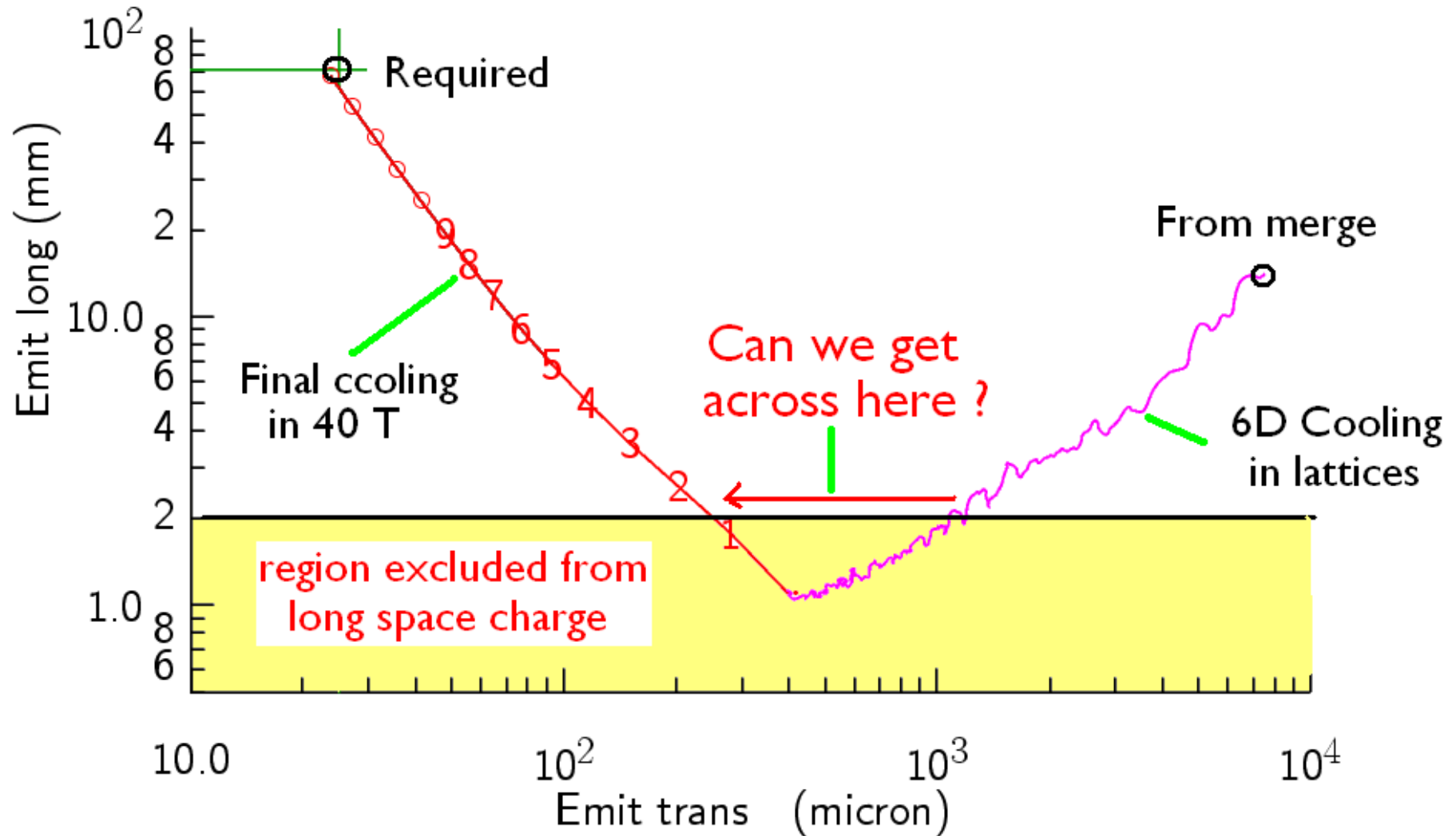
# Net focusing with rf for JLAB scheme



- With a Gaussian distribution the bunch center is de-focused !
- But since this condition is reached slowly, the bunch will tend towards the parabolic
- Simulation is required to see if this is acceptable
- But, for the moment we will assume that it is not acceptable

# The new cooling challenge

Probably the only way to reduce the Longitudinal space charge is to lower the longitudinal emittance, so the question is

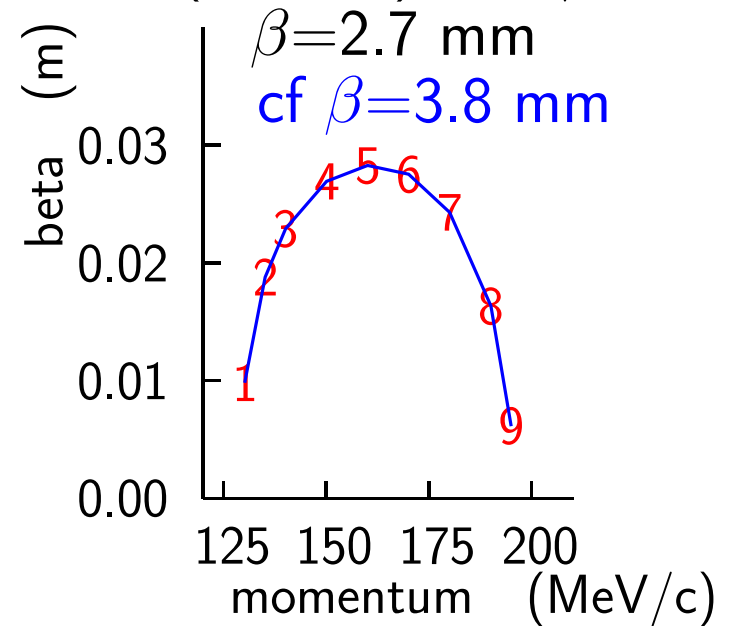
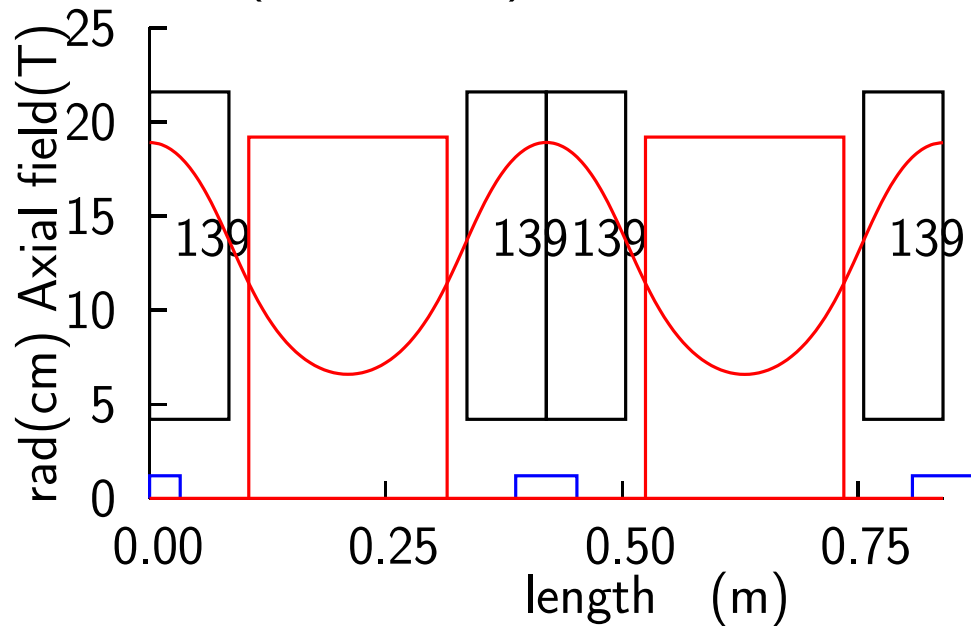


# Philosophy

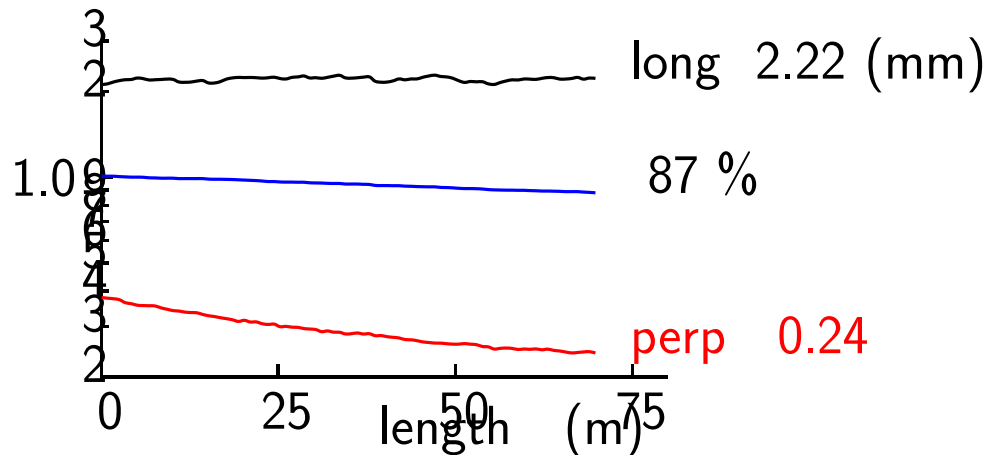
- Keep  $\epsilon_{\parallel} \geq 2$  mm (vs. 1 mm) to reduce space charge
- Weaken emittance exchange n tapered cooling after merge, to keep  $\epsilon_{\parallel}$  above 2 mm
- Design special stages for cooling only in transverse direction, but with enough exchange to keep long emit constant
  - Use non-flip (Fermi) cells to lower required fields and current densities
  - Lower momentum to further lower required fields and current densities
  - Reduce cell length, restoring fields and densities, lowering betas, and giving more absorbers per meter

# New Non-flip lattice allows 6D to lower $\epsilon_{\perp}$

- 42 (vs. 68.75) cm cell, momentum 160 (vs. 200) MeV/c

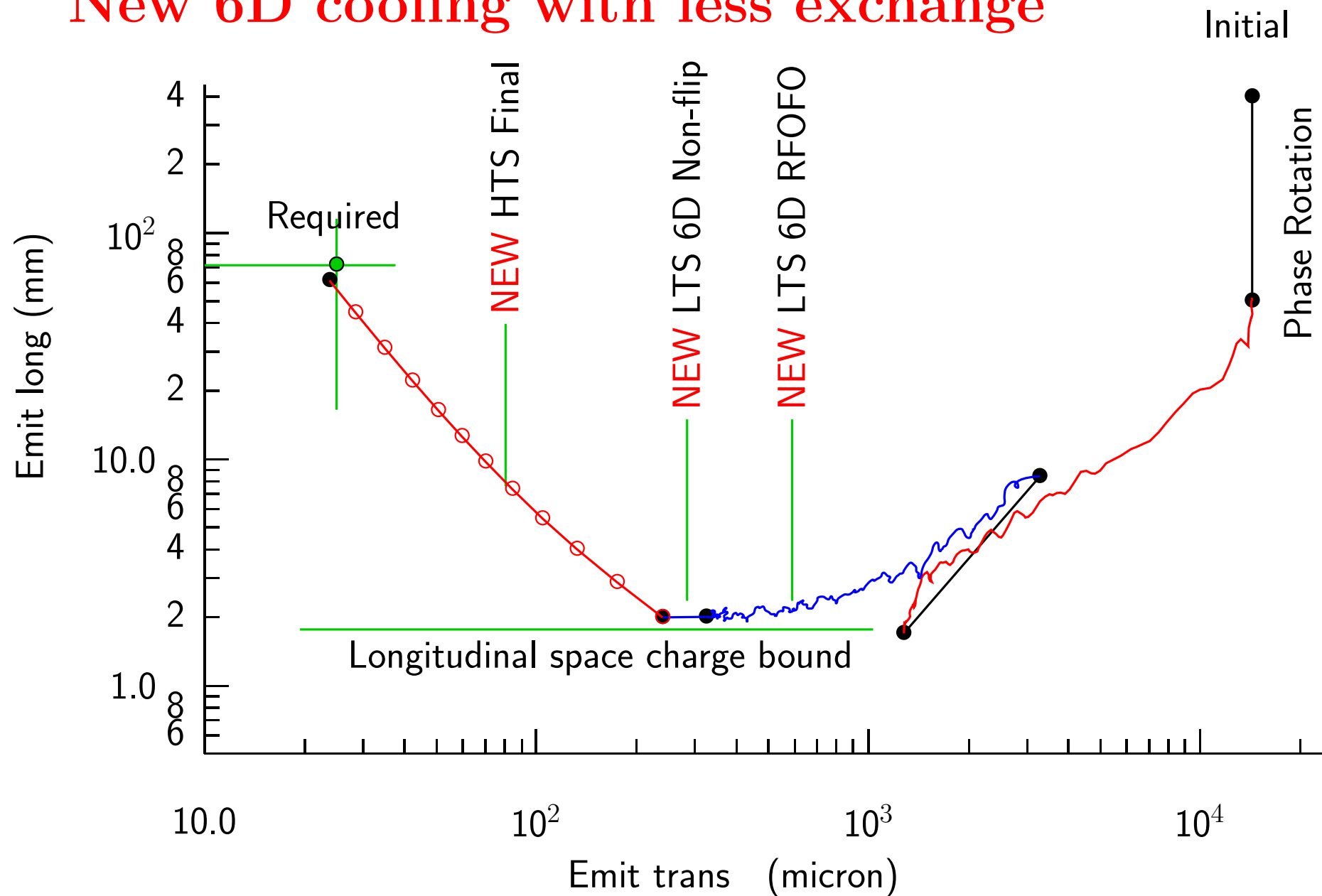


ICOOOL  
Simulation

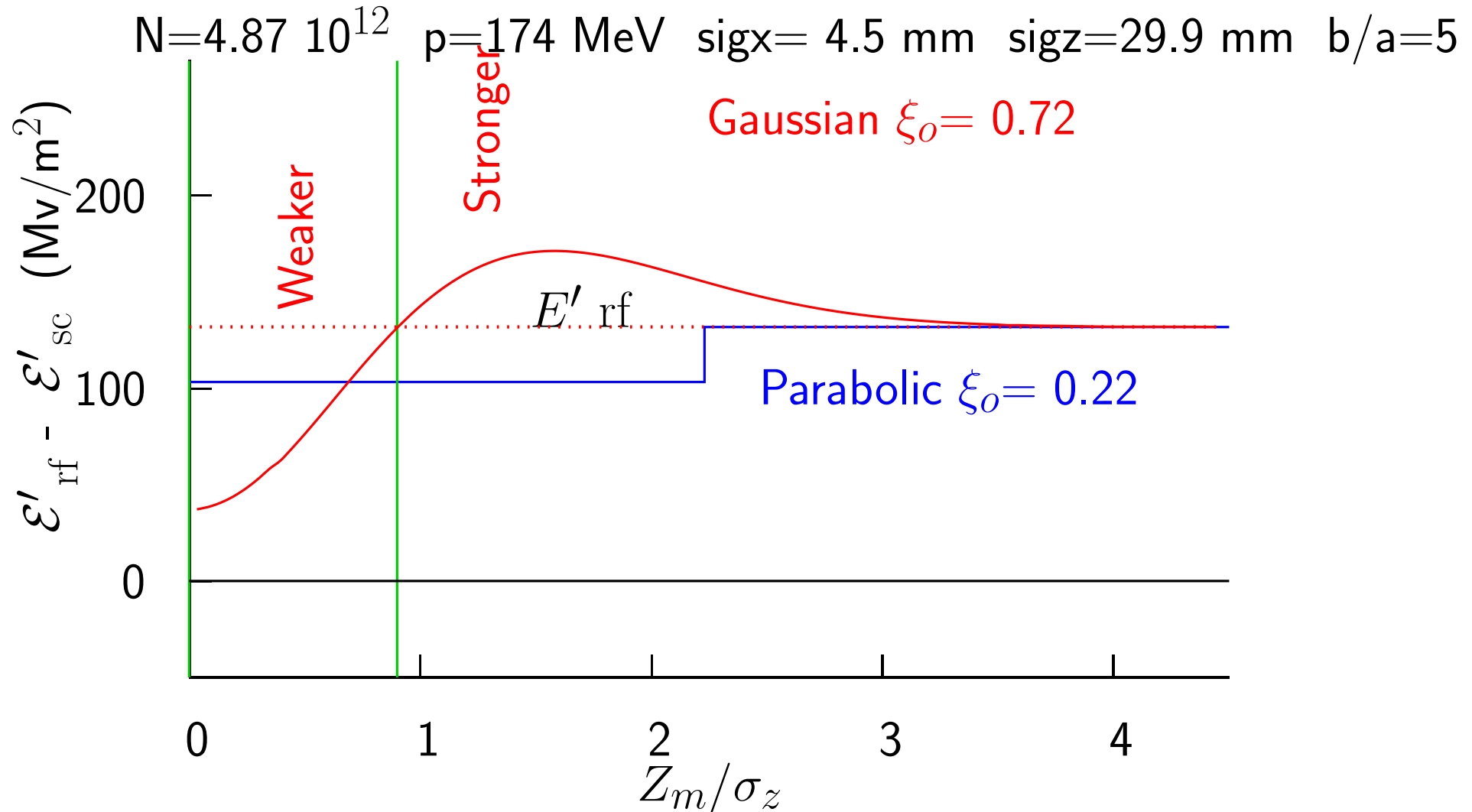


Note Brf=12 T cf 6 T for RFOFO

# New 6D cooling with less exchange



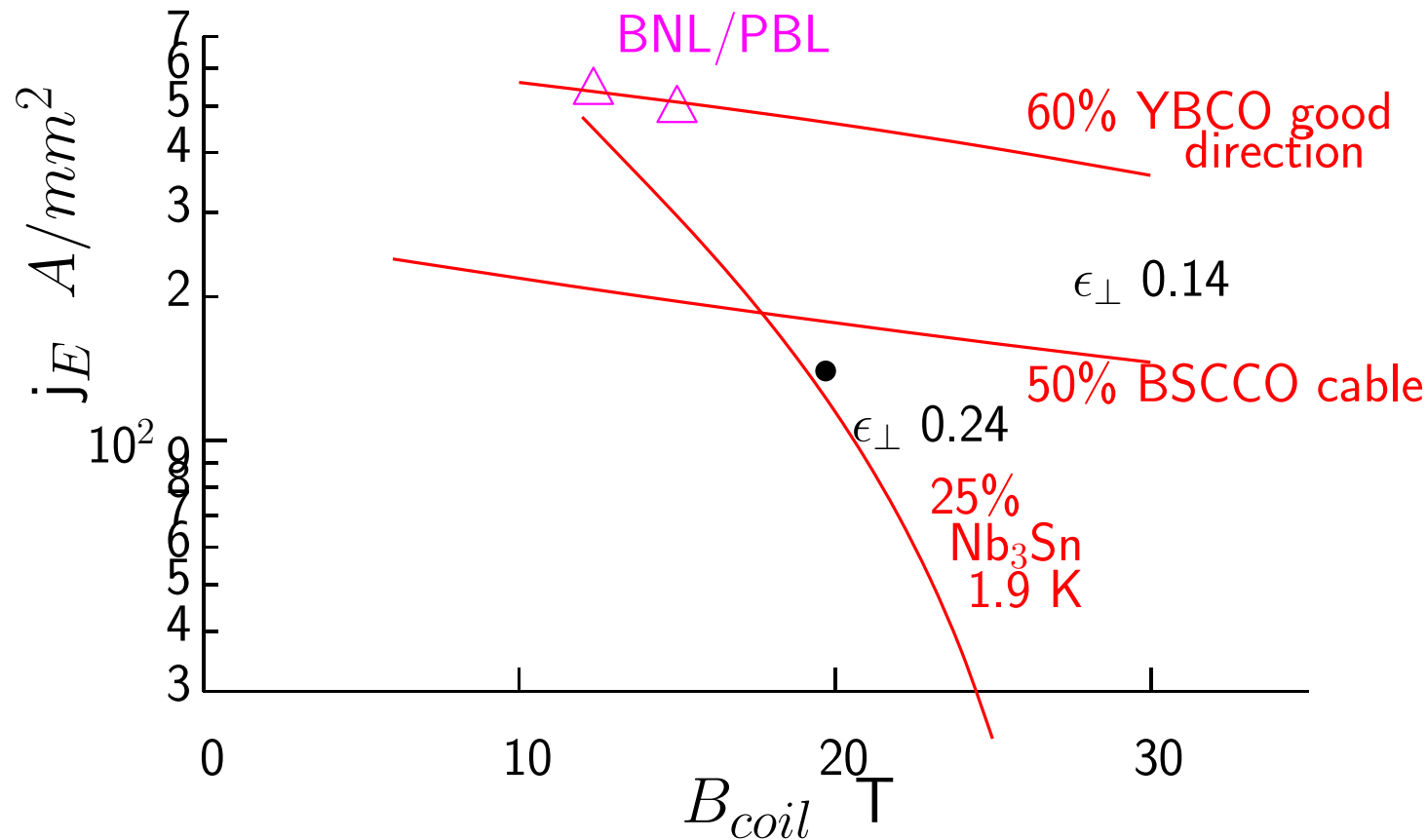
# Space Charge with new lattice



- Need rf increase from 20.0 MV/m  $\rightarrow$  23.9 MV/m
- Assumed acceptable



# Superconductor requirements and limits

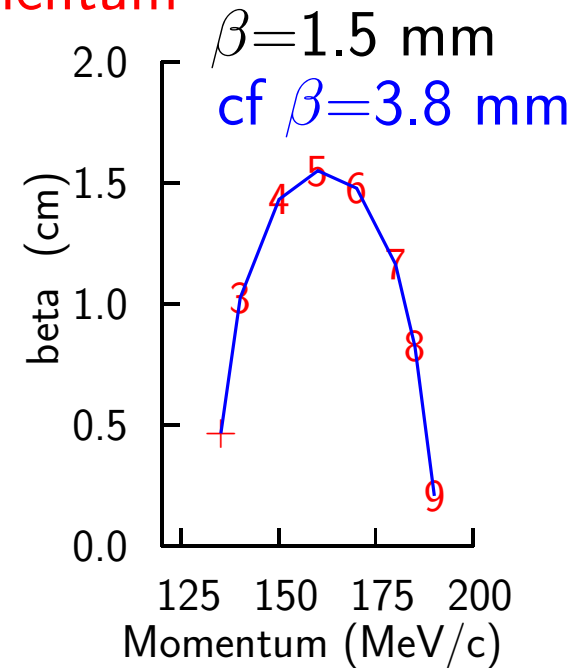
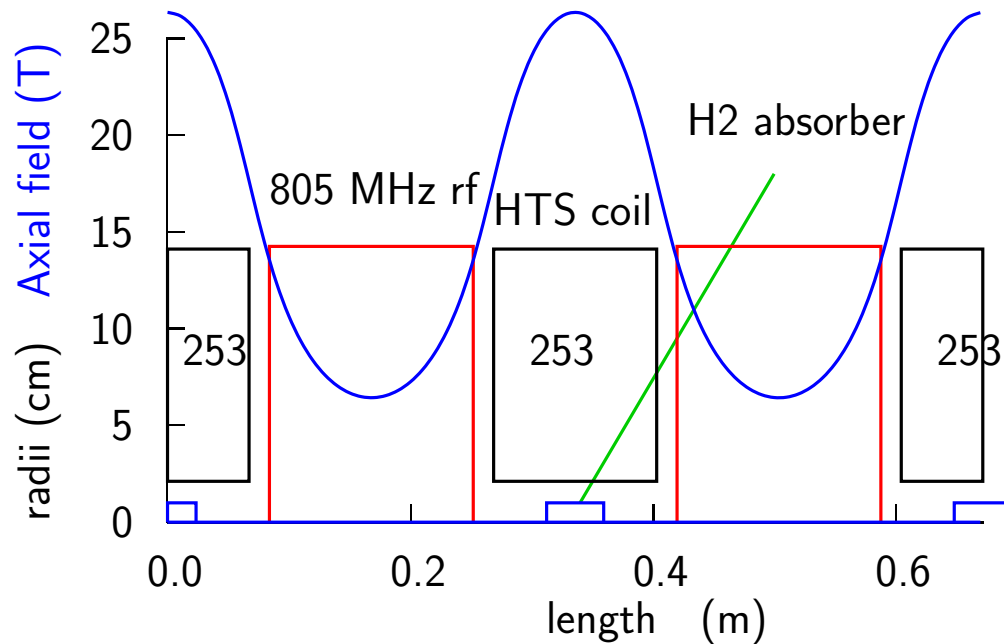


- 60% of YBCO tape justified by BNL/PBL test coils
- 25% of Nb<sub>3</sub>Sn for stabilizer & ss from Bob Weggel 15 T design
- 50% of BSCCO assumes ss support, but no additional stabilizer
- Conservative: they assume uniform density based on highest field

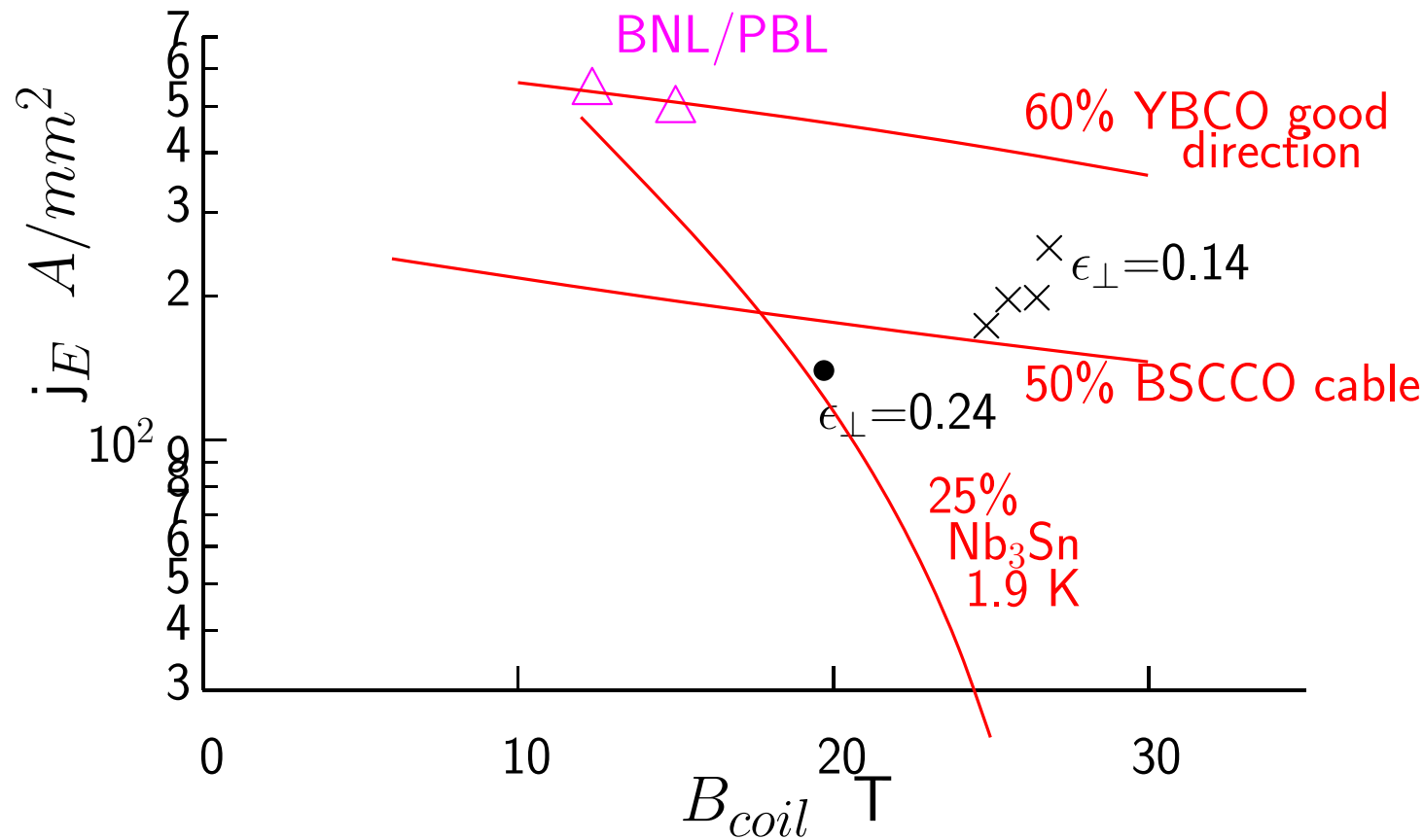
# Can we do better with HTS coils

	cell cm	Mom MeV/c	beta cm	emit mm	L cm	r1 cm	r2 cm	j A/mm <sup>2</sup>	Bo T	Bmax T
37h	41.0	200	2.8	0.24	16.8	4.2	21.6	174	23.6	24.7
38h	41.0	200	2.3	0.20	16.8	3.8	18.8	197	24.4	25.3
39h	41.0	200	1.9	0.17	16.8	2.6	17.6	199	26.0	26.2
40h	33.6	160	1.5	0.14	13.4	2.1	14.1	253	26.3	26.6

## Example of lattice 40h and $\beta$ vs. Momentum



# 6D cooling HTS requirements and limits

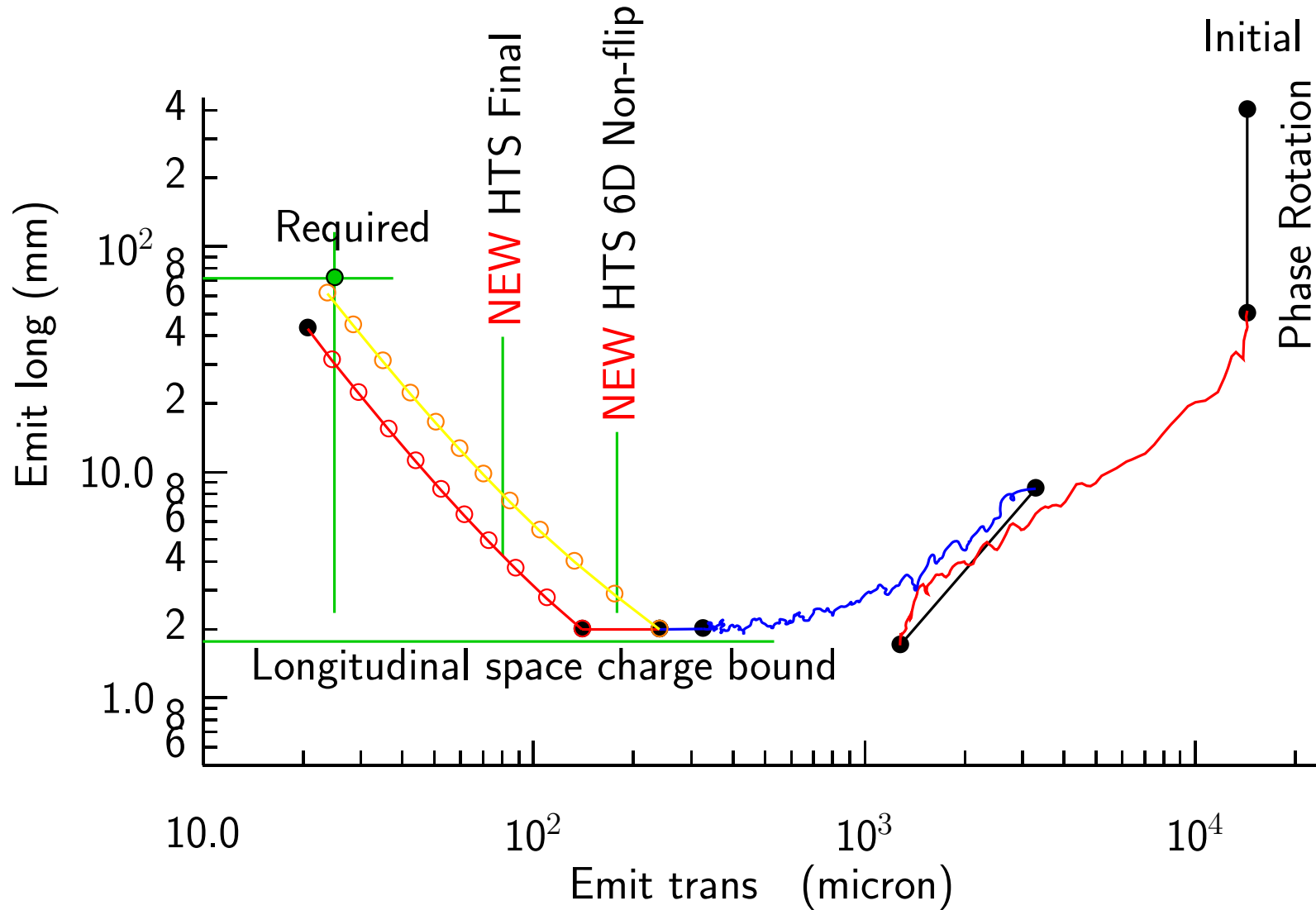


● earlier lattice using Nb<sub>3</sub>Sn

× later lattices using HTS

- Only YBCO is suitable for these 6D cooling lattices

# With HTS 6D and new Final Cooling



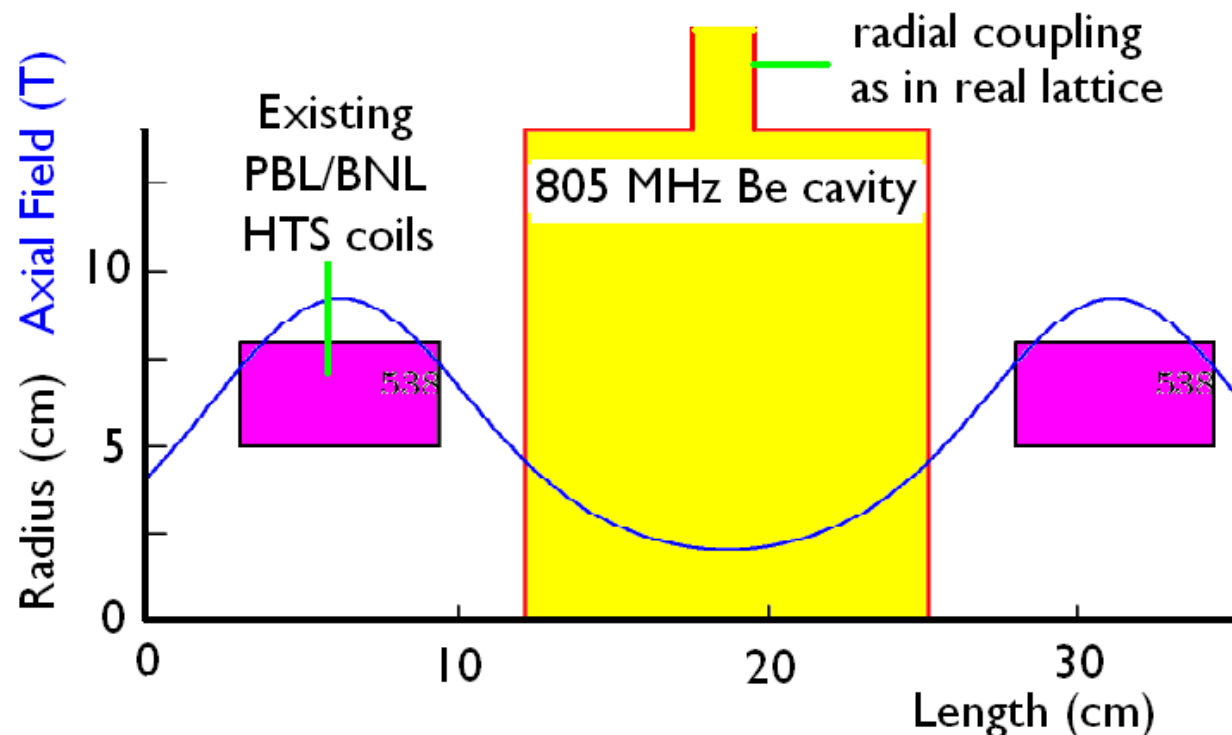
- $\epsilon_{\perp} = 20.7$  (cf. 25)  $\mu\text{m}$   $\epsilon_{\parallel} = 43$  (cf. 72) mm
- Allows more realistic allowances for dilution in acceleration

# Conclusion

- Allen et al have full simulations of the parabolic case
- We can express our more Gaussian bunches as sums of parabolic distributions
- Long. space charge for JLAB scheme is probably unacceptable
- A solution is to increase the minimum longitudinal emittance
- We do this by:
  - reducing emittance exchange in tapered lattice,
  - adding a new lattice with non-flip fields
  - but field on rf now  $\approx 12$  T (vs. 6 T)  
will discuss more on Wednesday
- Longitudinal Space Charge now looks acceptable
- With HTS 6D cooling we can get to lower final emittances giving more realistic allowed dilutions in acceleration

# Challenge testing rf in high magnetic fields

- How to get  $\approx 12$  T fields ?
- Use coils developed BBL/BNL coils for 40 T (Gupta)
- First exp uses existing YBCO coils
- Needs side coupled cavity, as discussed by Zenghai Li
- Both require rf at 77 degrees and cryostat



# Later experiment

- Also use Nb<sub>3</sub>Sn and NiTi coils
- Need Nb<sub>3</sub>Sn coils as under design for Phase I outsert of 35 T
- Geometry now near real
- Field on rf 10 T, Field max 18 T, both near real

